

XSPEC - Implementation of analytical models

Michal Dovčiak

Astronomical Institute
Academy of Sciences of the Czech Republic
Prague

X-ray data modelling of accreting black holes
Hotel Belaria, Hradec nad Moravicí, Czech Republic
17th – 25th July 2021

Introduction

Information available in XSPEC manual [Appendix C: Adding Models to XSPEC](https://heasarc.gsfc.nasa.gov/xanadu/xspec/manual/node313.html)
(<https://heasarc.gsfc.nasa.gov/xanadu/xspec/manual/node313.html>)

- ▶ Create subroutine with the new XSPEC model
- ▶ Create description file for the model and its parameters
- ▶ Compile the model inside the XSPEC (once)
- ▶ Load the model library
- ▶ Use the model

Example I – power-law with sharp energy cut-offs – the subroutine

```
#include <math.h>

void pwlc0(const double *ear, int ne, const double *param, int ifl,
           double *photar, double *photer, const char* init) {

    int    ie;
    double e1, e2;

    for (ie = 0; ie < ne; ie++){
        if ( ear[ie+1] <= param[1] ){
            photar[ie] = 0.;
            continue;
        }else if ( ear[ie] <= param[1] ) e1 = param[1];
        else e1 = ear[ie];
        if ( ear[ie] >= param[2] ){
            photar[ie] = 0.;
            continue;
        }else if ( ear[ie+1] >= param[2] ) e2 = param[2];
        else e2 = ear[ie+1];

        if(param[0] != 1.)
            photar[ie] = ( pow( e2, -param[0] + 1. ) - pow( e1, -param[0] + 1. ) )
                        / ( -param[0] + 1. );
        else photar[ie] = log( e2 / e1 );
    }

    return;
}
```

- ▶ **ear** – energy array with **ne+1** values for energy bin boundaries
- ▶ **ne** – size of flux array
- ▶ **param** – parameter values
- ▶ **ifl** – the spectrum number being calculated
- ▶ **photar** – photon flux array with **ne** values defined per energy bin, i.e. not per keV!
- ▶ **photer** – photon flux error array with **ne** values
- ▶ **init** – this string is read on initialization and available to the model during execution (optional)

Example I – power-law with sharp energy cut-offs – the description

Entry for a model description in `lmodel.dat` file:

```
pwlco      3  0.          1.0e20          c_pwlco  add  0  0
PhoIndex " "      2.0      1.2      1.2      3.      3.      0.1
E0      keV      0.1      0.001  0.001  1.      1.      0.1
Ecut     keV     100.      10.      10.      1000.   1000.   10.
```

The first line for each model:

- ▶ the model name
- ▶ the number of parameters
- ▶ the low and high energies for which the model is valid
- ▶ the name of the subroutine to be called
- ▶ the type of model (add, mul, mix, or con, or acn)
- ▶ the flag to be set to 1 if model variances are calculated by `modelfunc`
- ▶ the flag to be set to 1 if the model should be forced to perform a calculation for each spectrum

Example I – power-law with sharp energy cut-offs – the description

Entry for a model description in `lmodel.dat` file:

```
pwlco      3  0.          1.0e20          c_pwlco  add  0  0
PhoIndex " "      2.0      1.2      1.2      3.      3.      0.1
E0      keV      0.1      0.001  0.001  1.      1.      0.1
Ecut     keV     100.      10.      10.      1000.  1000.  10.
```

The remaining lines specify each parameter in the model:

- ▶ the parameter name
- ▶ optional units label
- ▶ the default parameter value
- ▶ hard min and soft min
- ▶ soft max and hard max
- ▶ fit delta

Example I – power-law with sharp energy cut-offs – the description

Entry for a model description in `lmodel.dat` file:

```
pwlco      3  0.          1.0e20          c_pwlco  add  0  0
PhoIndex "  "      2.0      1.2      1.2      3.      3.      0.1
E0        keV      0.1      0.001    0.001    1.      1.      0.1
Ecut      keV      100.     10.      10.      1000.    1000.    10.
```

Three special types of parameters can be used:

- ▶ **scale** parameter – its name is prefixed with a *, it cannot be made variable and may be linked only to another scale parameter, only the initial value is given
- ▶ **switch** parameter – its name is prefixed with a \$, it is not used directly but switches the model's mode of operation, it only has 2 fields: its name and an integer value
- ▶ **periodic** parameter – **P** is added at the end of the line, if it tries to exceed its hard limits during the fit, it will be assigned a value within its limits: $f(\max + \delta) = f(\min + \delta)$, $f(\min - \delta) = f(\max - \delta)$

Example II – model subroutine using FITS tables

```
#define REFSPECTRA "myrefl.fits\0"
#define NPAR 3

extern void tabint(float* ear, int ne, float* param, int npar,
                  const char* filenm, int ifl, const char* tabtyp,
                  float* photar, float* photer);

void myrefl(const double *ear, int ne, const double *param, int ifl,
            double *photar, double *photer, const char* init) {

char    refspectra[12] = REFSPECTRA;
int     ie;
float   fl_ear[ne+1], fl_photar[ne], fl_photer[ne];
float   fl_param[NPAR]={((float) param[0], (float) param[1], (float) param[2]);
const char* tabtyp="add";
double  flux;

//Let's read and interpolate the FITS tables using internal XSPEC routine tabint
//Note that we do not use errors here
for(ie = 0; ie <= ne; ie++) fl_ear[ie] = (float) ear[ie];
tabint(fl_ear, ne, fl_param, NPAR, refspectra, ifl, tabtyp, fl_photar, fl_photer);

//Let's e.g. renormalize to the total counts
flux=0.;
for (ie = 0; ie < ne; ie++) flux += fl_photar[ie];
for (ie = 0; ie < ne; ie++) photar[ie] = fl_photar[ie] / flux;

return;
}
```

How to install and use local models

Installation in XSPEC:

- ▶ the code is compiled inside XSPEC with the following command:
initpackage mypackage lmodel.dat /path/to/mypackage

To use local models inside XSPEC:

- ▶ the package needs to be loaded:
lmod mypackage /path/to/mypackage
- ▶ then any model from the package may be used, e.g.:
mo pwlco